The Kepler Mission: Are There Any Good Worlds Out There?







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31 December 2010



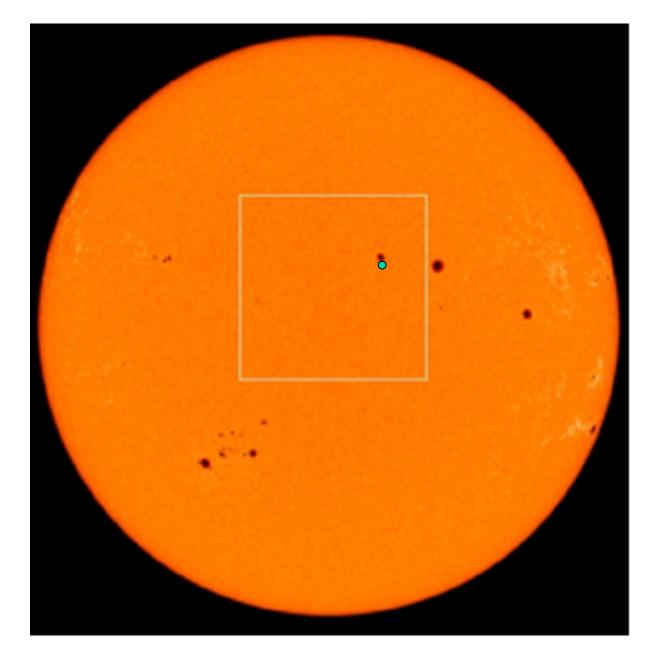
Overview

- Why Kepler?
- The Kepler Mission
- Processing Kepler Data
- Kepler Light Curves
 - New Discoveries
 - How is *Kepler* Changing the Field of Exoplanets



The Sun In Visible Light





Earths are much easier to find when UV light is blocked.

Earth-size planets Are about the size of a star spot.

Rapid motion and uniform repetition is used to distinguish planets from spots.





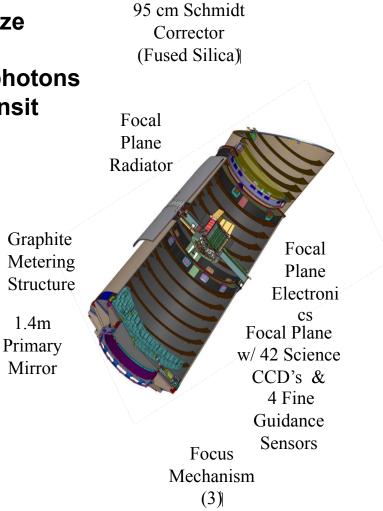
KEPLER: A Wide Field-of-View Photometer that Monitors 100,000 Stars for 3.5 yrs with Enough Precision to Find Earth-size Planets in the Habitable Zone

Use transit photometry to detect Earth-size planets

- 0.95 meter aperture provides enough photons
- Observe for several years to detect transit patterns
- Monitor a single large area on the sky continuously to avoid missing transits
- Use heliocentric orbit

Get statistically valid results by monitoring 100,000 stars

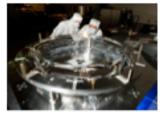
- Wide Field-of-view telescope
- Large array of CCD detectors





Kepler Spacecraft

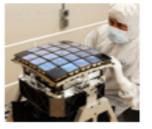




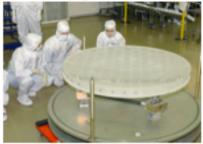
Schmidt Corrector 0.95 m dia.



Spider with Focal Plane and Local Detector Electronics



Focal Plane 95 Mega pixels, 42 CCDs



12/30/10Primary Mirror 1.4 m dia., 85% lt. wt.



Fully assembled Kepler photometer Mounted on the spacecraft



Sunshade



Upper Telescope Housing



Lower Telescope Housing



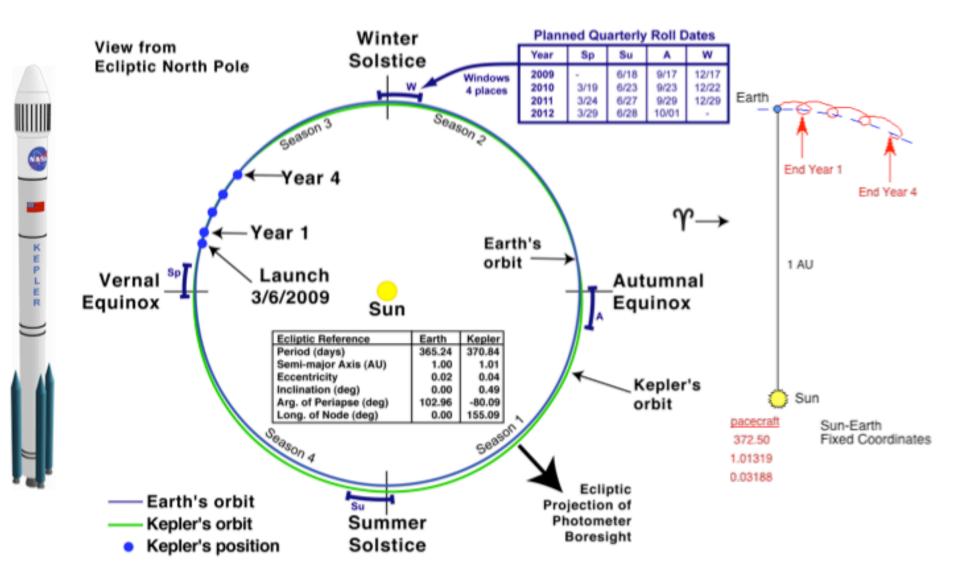
Spacecraft bus integration

6



Earth-Trailing Heliocentric Orbit



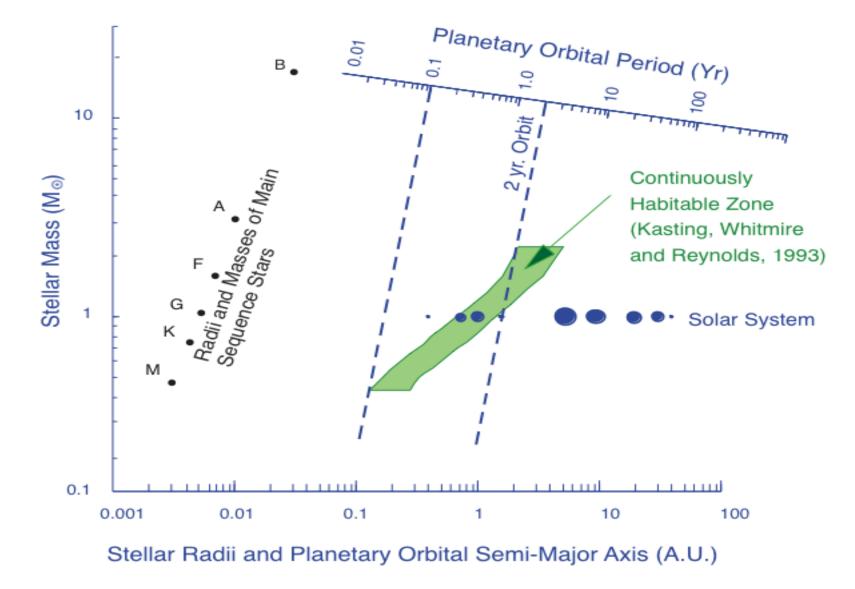




The Habitable Zone for Various Stellar Spectral Types



A Search for Habitable Planets

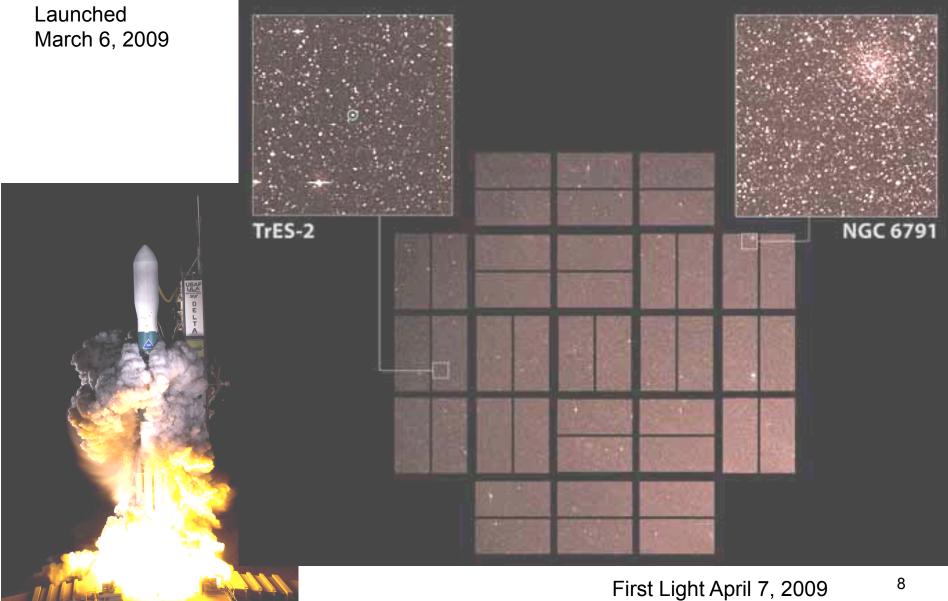


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First Light Full Field Image

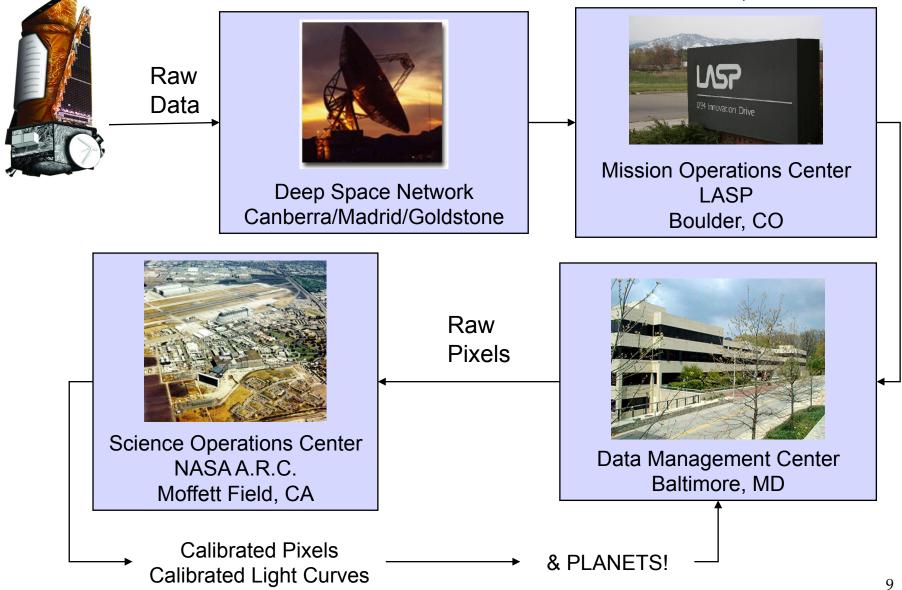






Kepler Mission Data Flow

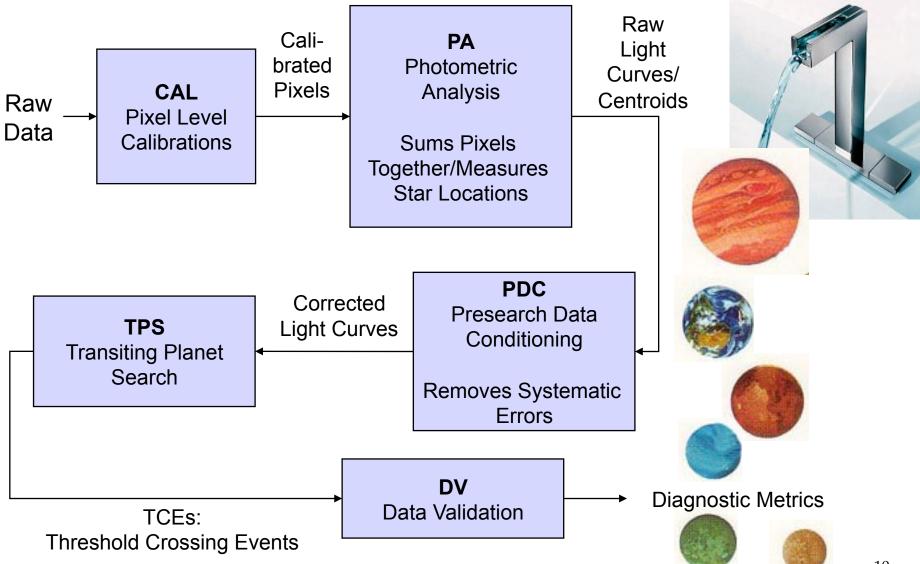






The SOC Pipeline: From Pixels To Planets

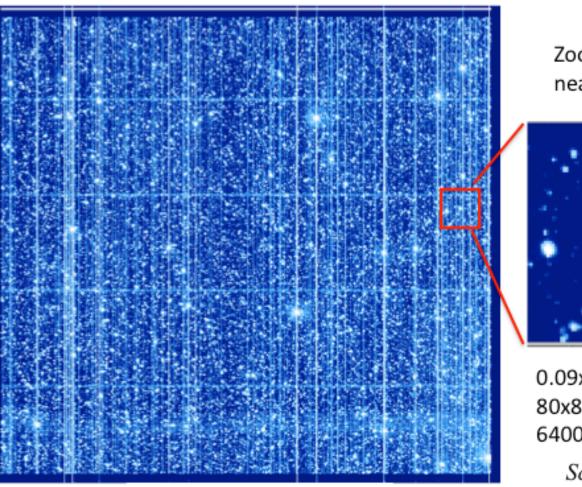






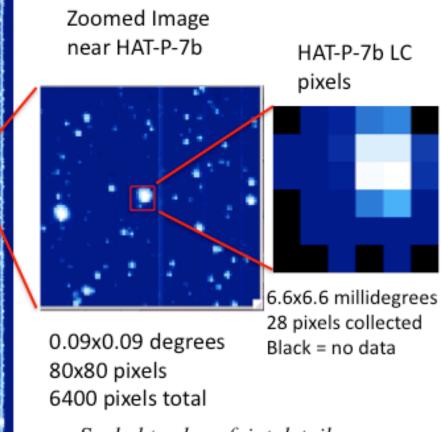


A Search for Habitable Planets



Module 17 Output 2

1.13 (h) x1.22 (w) degrees

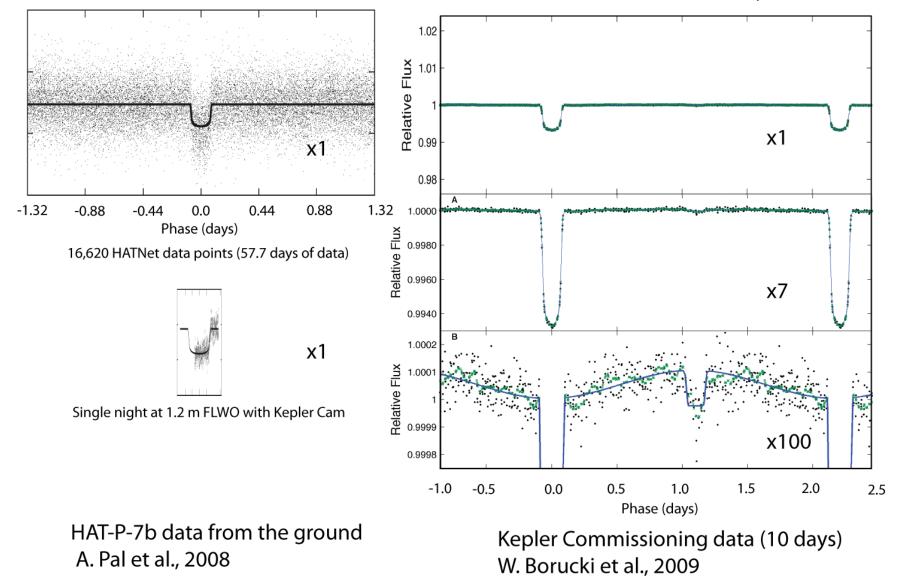


Scaled to show faint detail



HAT-P-7b Ground vs. Space

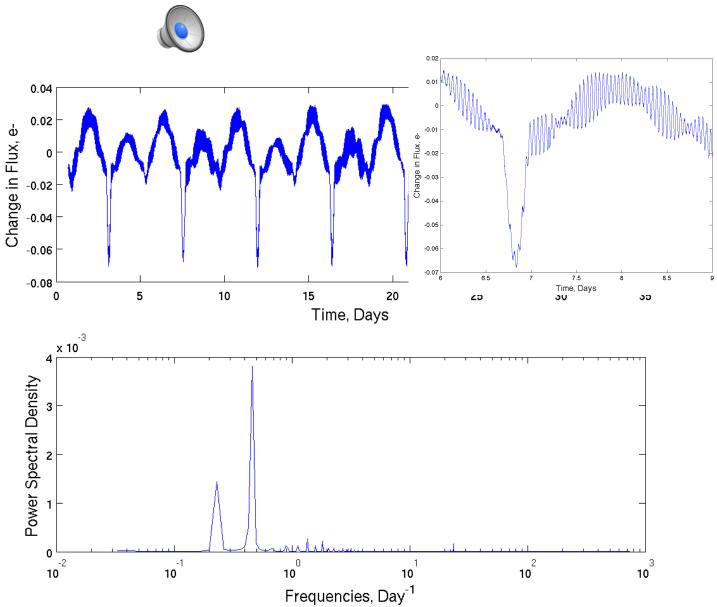






Music From the Stars (3)







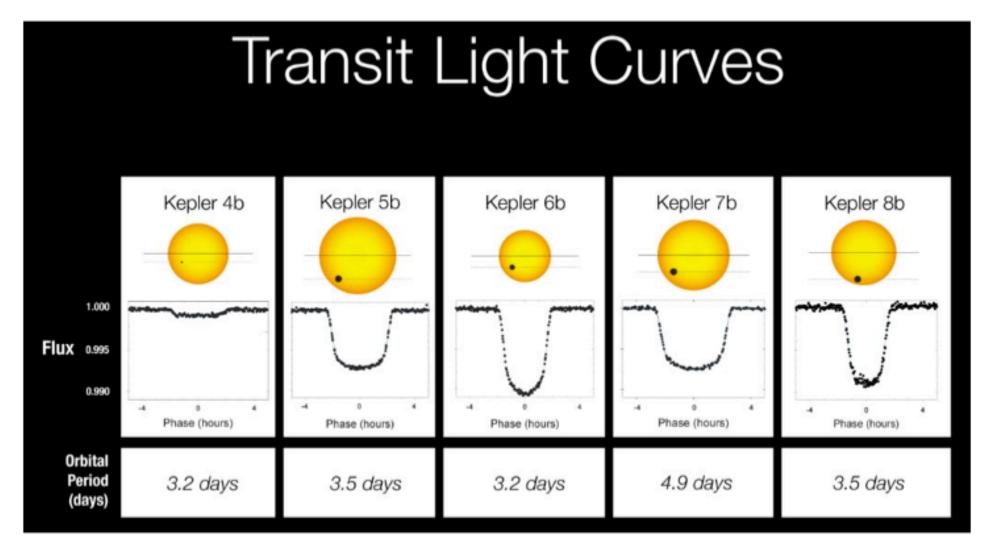
Music From the Stars (4)



A Search for Habitable Planets 0.08 0.06 0.1 0.04 Change in Flux, e-Change in Flux, e-0.05 0 -0.02 -0.04 -0.05 -0.06 — 0 1.8 0.2 0.4 0.6 0.8 1.2 1.4 1.6 2 1 Time, Days -0.1 10 15 20 0 5 25 30 35 Time, Days 0.012 Power Spectral Density = 10.0 Bootral Density = 200.0 Control Density 0 , ∟ -2 10 -1 10 10⁰ 10² 3 10³ 10¹ Frequencies, Day⁻¹



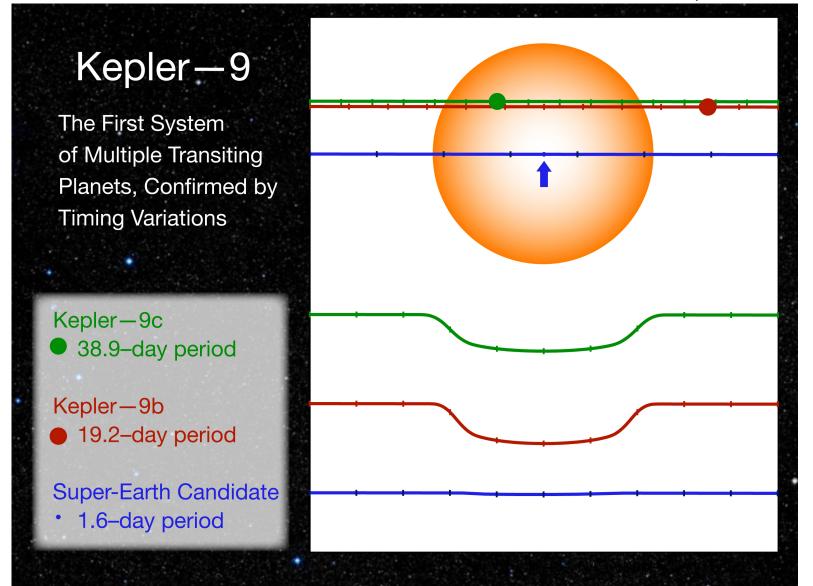






Keple-9b, c and d

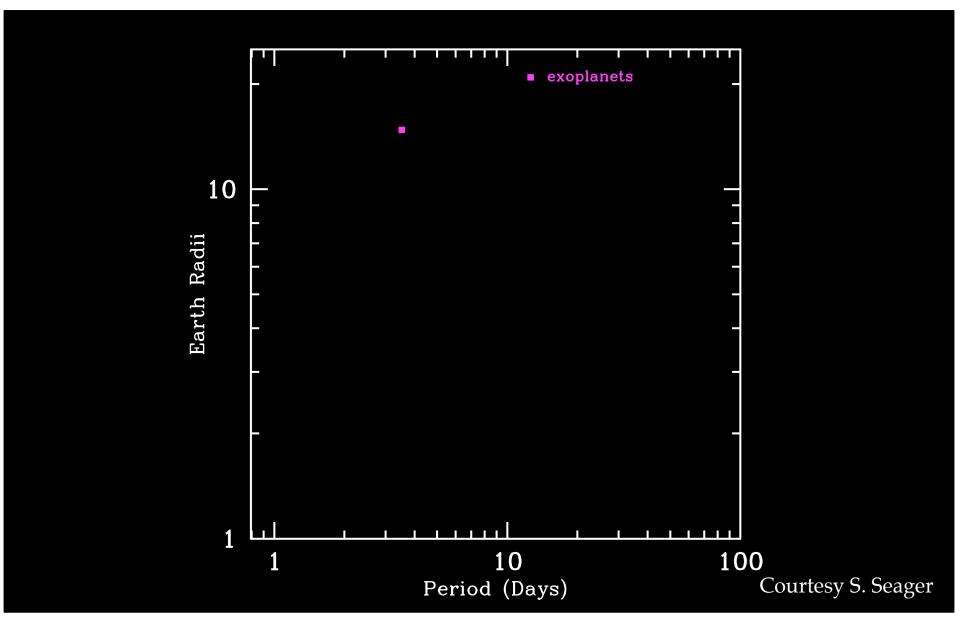






Transiting Planets 2000

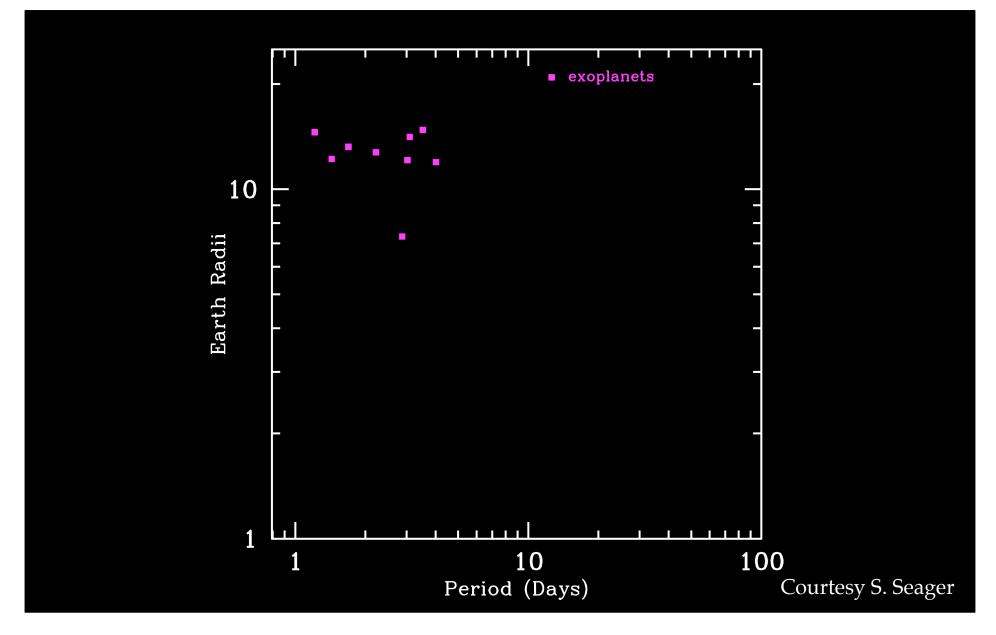






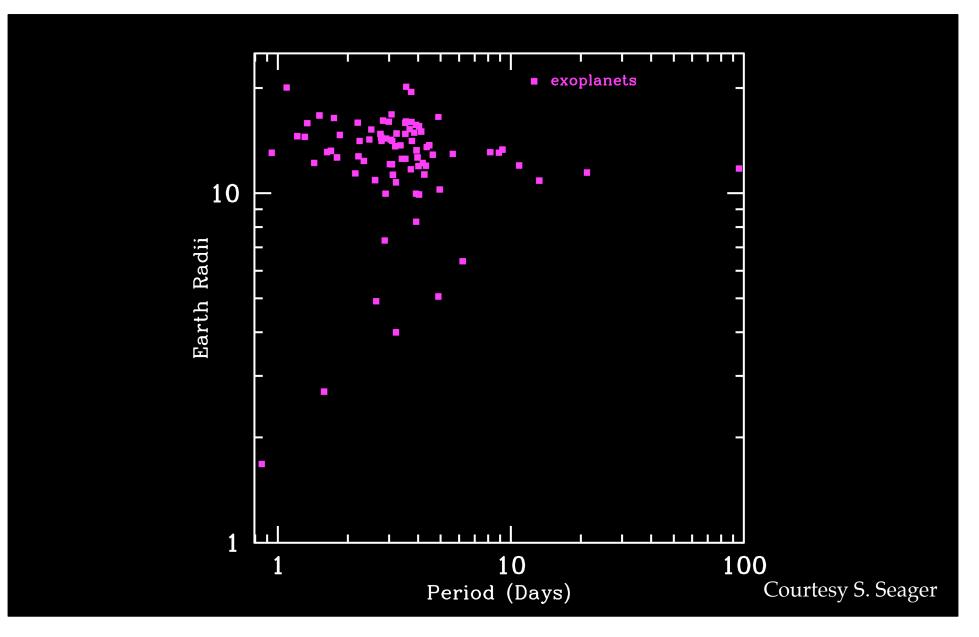
Transiting Planets 2005





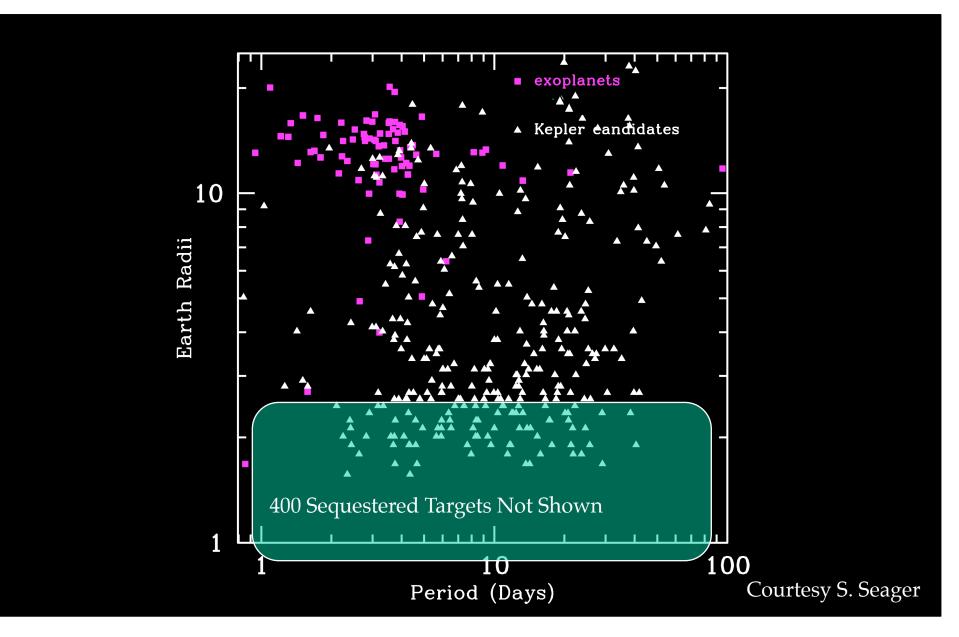














How is Kepler Changing the Field Of Exoplanets?



A Search for Habitable Planets



"Well, this mission answers at least <u>one</u> big question: Are there other planets like ours in the universe?"

... the ways by which men arrive at knowledge of the celestial things are hardly less wonderful than the nature of these things themselves.

— Johannes Kepler 1571-1630

Do there exist many worlds or is there but a single one? This is one of the most noble and exalted questions in the study of Nature

— Saint Albertus Magnus 1206-1280 Scholar, Patron Saint of Scientists